

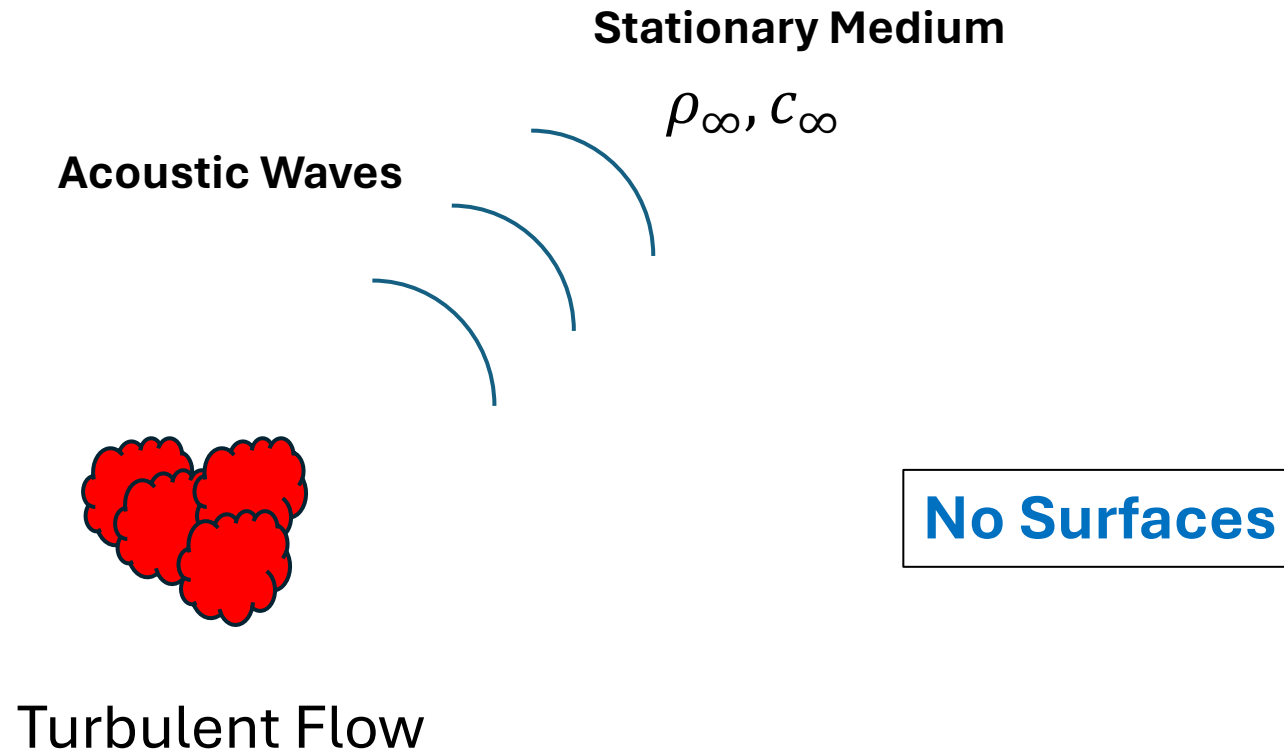
Lighthill's Acoustic Analogy

Lighthill's Acoustic Analogy Module

Outline of Topics

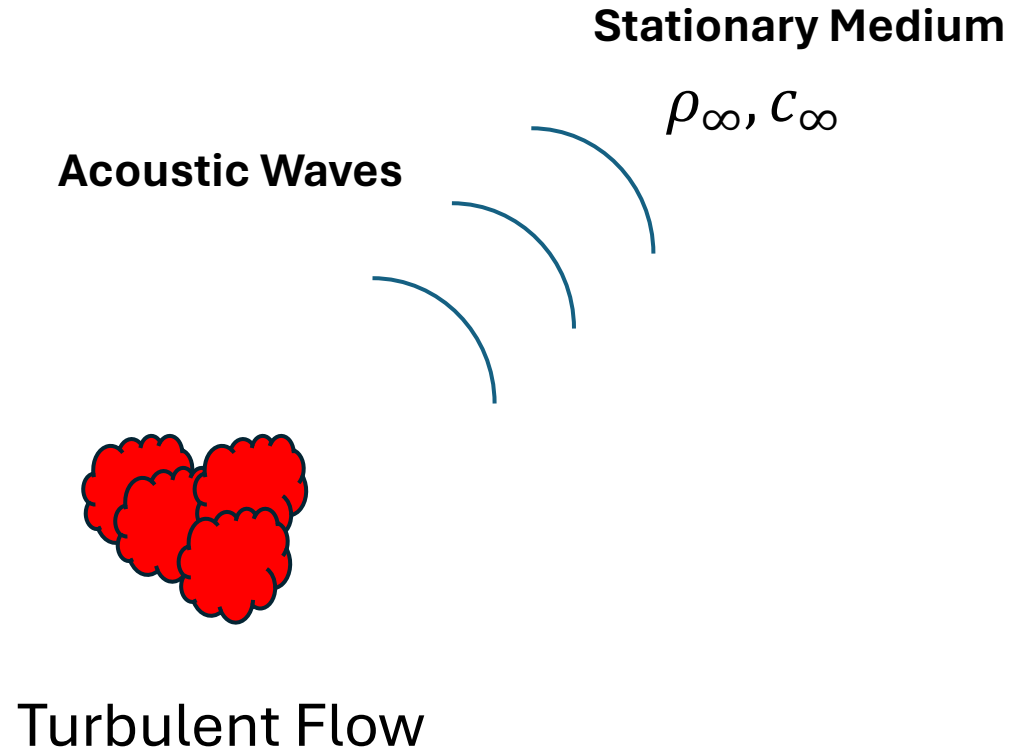
- The Concept $NS=0$
- Lighthill's Wave Equation
- The Free Field Solutions
- Limitations
- Near and Far Field Sound, U^8
- Surfaces, the FW-H equation
- Moving Sources
- Dipole, Thickness and Edge Noise
- Scaling on Flow Speed

Lighthill's Analogy



What is the sound level in the stationary medium?

Lighthill's Analogy



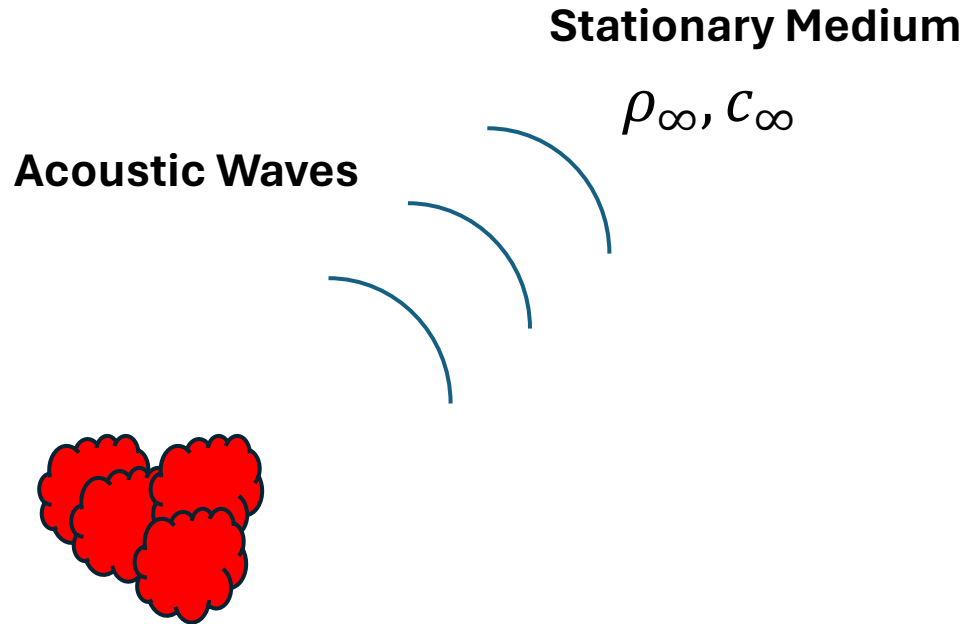
***The Navier Stokes
Equations Appy
Everywhere***

Momentum $NS=0$

Continuity $\frac{\partial M}{\partial t} = 0$

$$\nabla \cdot NS - \frac{\partial^2 M}{\partial t^2} = 0$$

Lighthill's Analogy



In the Stationary medium

$$\frac{\partial^2 \rho'}{\partial t^2} - c_{\infty}^2 \nabla^2 \rho' = 0$$

Adding and subtracting this from

$$\nabla \cdot \mathbf{NS} - \frac{\partial^2 M}{\partial t^2} = 0$$

$$\frac{\partial^2 \rho'}{\partial t^2} - c_{\infty}^2 \nabla^2 \rho' + \left\{ \nabla \cdot \mathbf{NS} - \frac{\partial^2 M}{\partial t^2} - \left(\frac{\partial^2 \rho'}{\partial t^2} - c_{\infty}^2 \nabla^2 \rho' \right) \right\} = 0$$

Lighthill's Analogy

Simplifying Terms

Stationary Medium

ρ_∞, c_∞

Acoustic Waves



$$\left\{ \nabla \cdot \mathbf{NS} - \frac{\partial^2 M}{\partial t^2} - \left(\frac{\partial^2 \rho'}{\partial t^2} - c_\infty^2 \nabla^2 \rho' \right) \right\} = \frac{-\partial^2 T_{ij}}{\partial x_i \partial x_j}$$

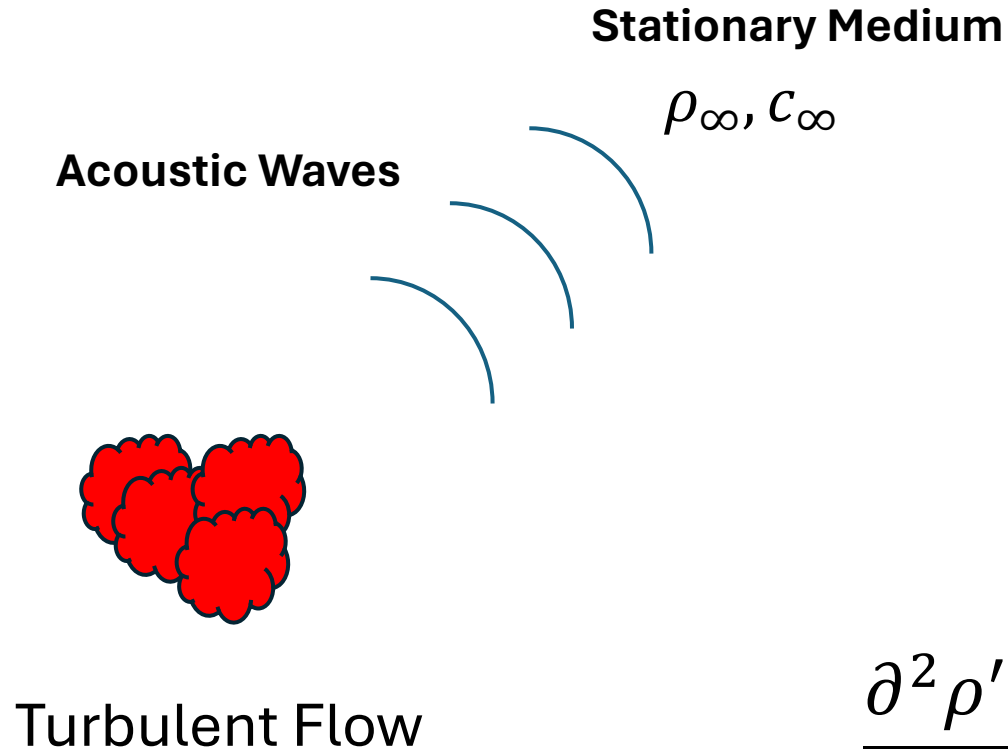
$$T_{ij} = \rho v_i v_j + (p_{ij} - \rho' c_\infty^2 \delta_{ij})$$

Turbulent Flow

$$\frac{\partial^2 \rho'}{\partial t^2} - c_\infty^2 \nabla^2 \rho' - \frac{\partial^2 T_{ij}}{\partial x_i \partial x_j} = 0$$

Zero in far field
stationary medium

Lighthill's Analogy

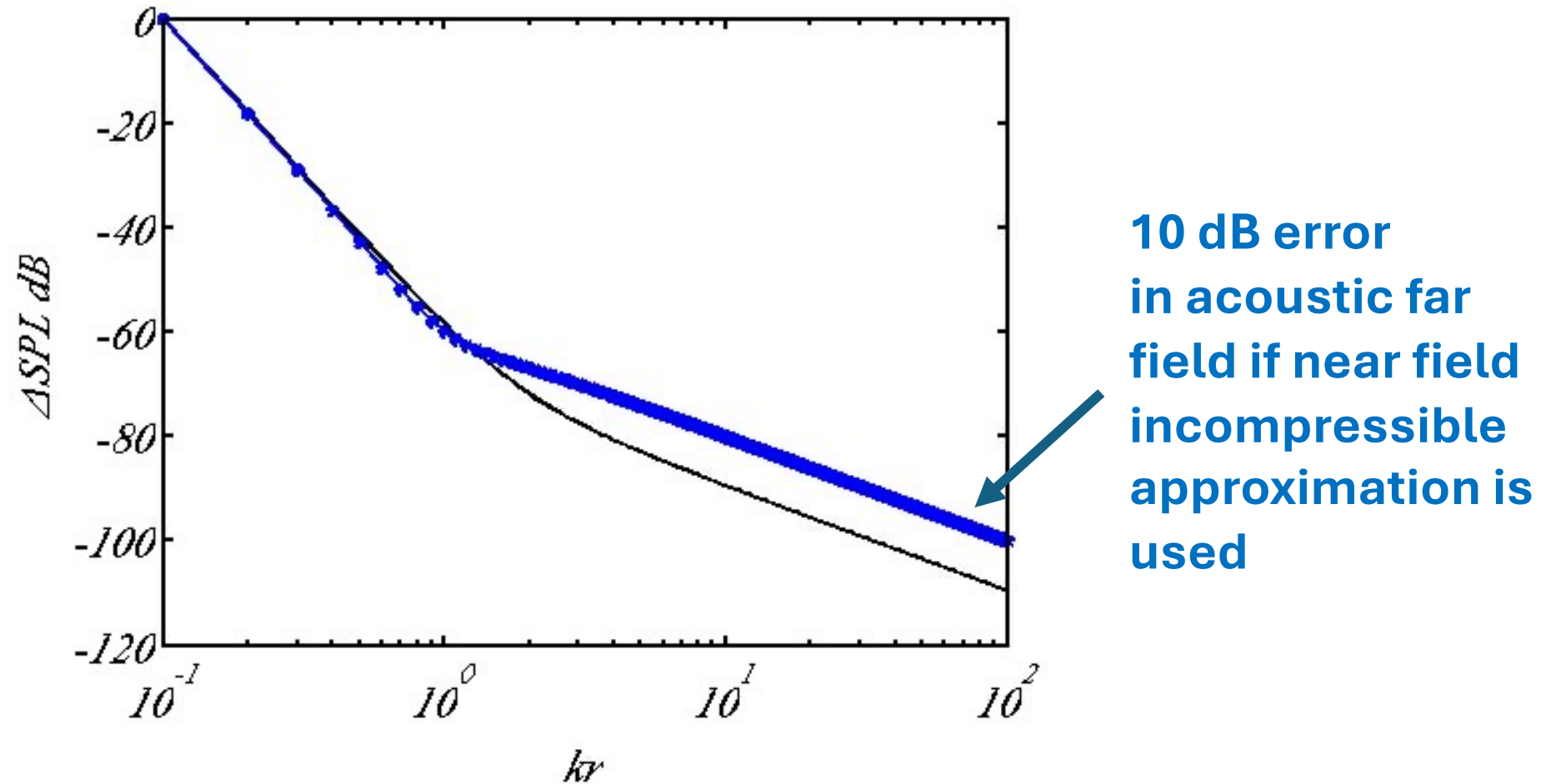


Lighthill's equation is an exact re-arrangement of the equations of motion that asymptotes to the wave equation in the far field...it is an interpolation method!

$$\frac{\partial^2 \rho'}{\partial t^2} - c_\infty^2 \nabla^2 \rho' - \frac{\partial^2 T_{ij}}{\partial x_i \partial x_j} = 0$$

Tends to zero in far field
stationary medium

The Effect of the incompressible source approximation at 45 deg. for longitudinal quadrupole



Lighthill's equation eliminates this type of error (in principle)

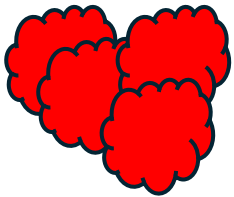
Lighthill's Analogy

Lighthill's Stress Tensor

Stationary Medium

$$\rho_{\infty}, c_{\infty}$$

Acoustic Waves



Turbulent Flow

$$T_{ij} = \rho v_i v_j + (p_{ij} - \rho' c_{\infty}^2 \delta_{ij})$$

$$\rho = \rho_{\infty} + \rho'$$

$$v_i = U_i + u_i$$

$$p_{ij} = p \delta_{ij} - \sigma_{ij}$$

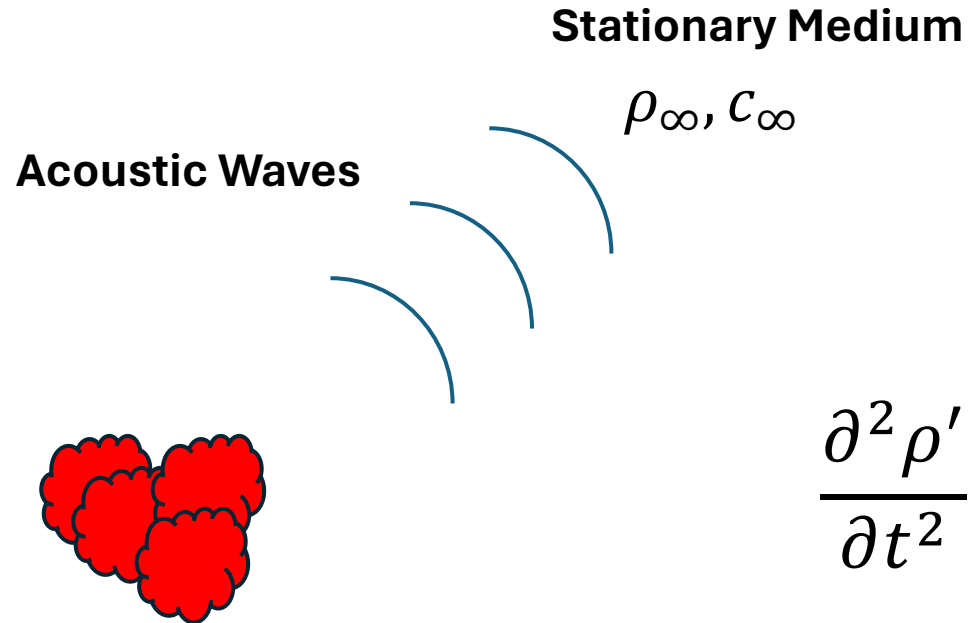
Mean shear Refraction

$$\begin{aligned} \rho v_i v_j = & \rho_{\infty} U_i U_j + \rho' U_i U_j \\ & + \rho (U_i u_j + u_i U_j) + \rho_{\infty} u_i u_j + \rho' u_i u_j \end{aligned}$$

Linearly coupled to pressure, so not independent of acoustic wave

Lighthill's Analogy

Lighthill's Stress Tensor



$$\frac{\partial^2 \rho'}{\partial t^2} - c_{\infty}^2 \nabla^2 \rho' - F(\rho') = \frac{\partial^2 T_{ij}}{\partial x_i \partial x_j} - F(\rho')$$

No exact analytical solutions to this equation

Greens Functions

$$\frac{\partial^2 \rho'}{\partial t^2} - c_\infty^2 \nabla^2 \rho' = \frac{\partial^2 T_{ij}}{\partial x_i \partial x_j}$$

Lighthill's equation is in the form of an inhomogeneous wave equation which can be solved using Greens functions

$$\frac{1}{c_\infty^2} \frac{\partial^2 \phi}{\partial \tau^2} - \nabla^2 \phi = Q(\mathbf{y}, \tau)$$

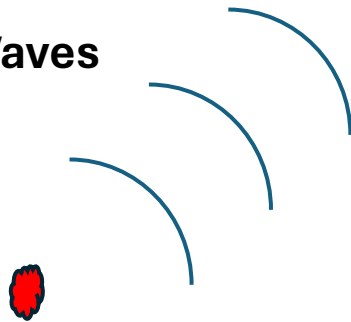
$$\phi(\mathbf{x}, t) = \int_V \int_{-T}^T Q(\mathbf{y}, \tau) G(\mathbf{x}, t | \mathbf{y}, \tau) dV d\tau$$

Greens Functions

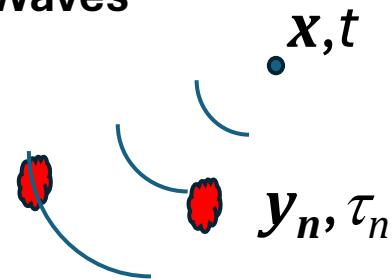
$$\phi(\mathbf{x}, t) = \int_V \int_{-T}^T Q(\mathbf{y}, \tau) G(\mathbf{x}, t | \mathbf{y}, \tau) dV d\tau$$

$$G(\mathbf{x}, t | \mathbf{y}, \tau) = \delta\left(t - \tau - \frac{|\mathbf{x} - \mathbf{y}|}{c_\infty}\right) / 4\pi |\mathbf{x} - \mathbf{y}|$$

Acoustic Waves



Acoustic Waves



The Greens function can be thought of as a filter function which weights the contribution of each source element and adds them up at the correct retarded time

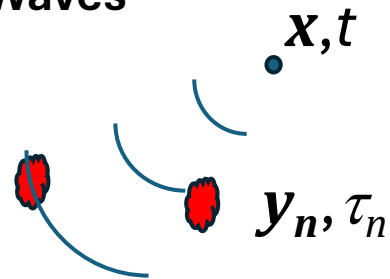
Greens Functions

$$\frac{1}{c_\infty^2} \frac{\partial^2 \phi}{\partial \tau^2} - \nabla^2 \phi = Q(\mathbf{y}, \tau)$$

$$\phi(\mathbf{x}, t) = \int_V \left[\frac{Q(\mathbf{y}, \tau)}{4\pi |\mathbf{x} - \mathbf{y}|} \right]_{\tau=\tau^*} dV$$

retarded time $\tau^* = t - \frac{|\mathbf{x} - \mathbf{y}|}{c_\infty}$

Acoustic Waves

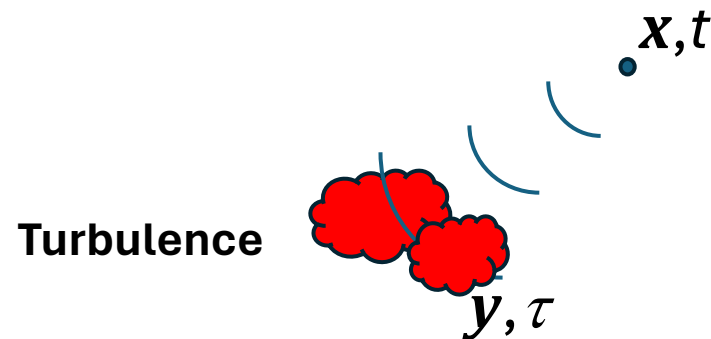


Lighthill's Equation

$$\frac{\partial^2 \rho'}{\partial t^2} - c_\infty^2 \nabla^2 \rho' = \frac{\partial^2 T_{ij}}{\partial x_i \partial x_j}$$

$$\rho' c_\infty^2(\mathbf{x}, t) = \int_V \frac{1}{4\pi |\mathbf{x} - \mathbf{y}|} \left[\frac{\partial^2 T_{ij}(\mathbf{y}, \tau)}{\partial y_i \partial y_j} \right]_{\tau=\tau^*} dV$$

$$\tau^* = t - \frac{|\mathbf{x} - \mathbf{y}|}{c_\infty}$$



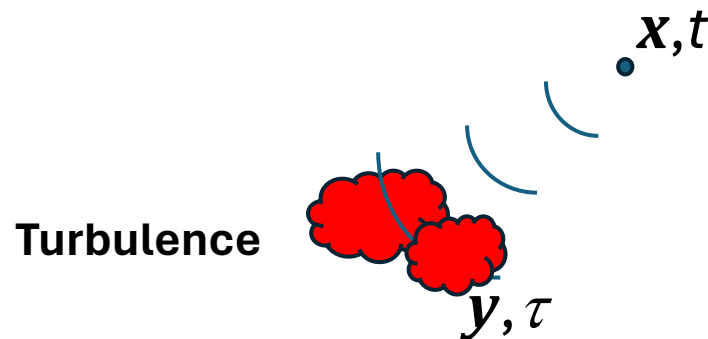
The sound in the stationary medium is the volume integral of the righthand side of Lighthill's wave equation

Lighthill's Equation for far field and small source regions $|\mathbf{x}| \gg |\mathbf{y}|, V \ll \lambda^3$

$$\rho' c_\infty^2(\mathbf{x}, t) = \frac{1}{4\pi|\mathbf{x}|} \left[\int_V \frac{\partial^2 T_{ij}(\mathbf{y}, \tau)}{\partial y_i \partial y_j} dV \right]_{\tau=\tau^*}$$

$$\tau^* \approx t - \frac{|\mathbf{x}|}{c_\infty}$$

This is a volume integral of a divergence so



$$\rho' c_\infty^2(\mathbf{x}, t) = \frac{1}{4\pi|\mathbf{x}|} \left[\int_{S_\infty} \frac{\partial T_{ij}(\mathbf{y}, \tau)}{\partial y_i} n_j dS \right]_{\tau=\tau^*} = 0$$

Use Ffowcs-Williams Observation

Use a modified variable

$$\frac{\partial^2 \rho'}{\partial t^2} - c_\infty^2 \nabla^2 \rho' = \frac{\partial^2 T_{ij}}{\partial x_i \partial x_j}$$

$$\frac{1}{c_\infty^2} \frac{\partial^2 \phi_{ij}}{\partial \tau^2} - \nabla^2 \phi_{ij} = T_{ij}(\mathbf{y}, \tau)$$

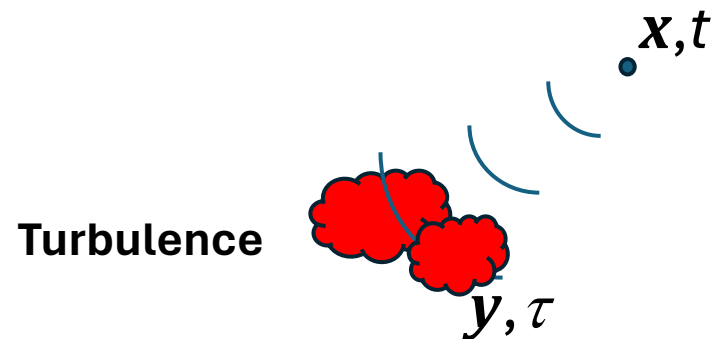
$$\phi_{ij}(\mathbf{x}, t) = \int_V \int_{-T}^T T_{ij}(\mathbf{y}, \tau) G(\mathbf{x}, t | \mathbf{y}, \tau) dV d\tau$$

$$\rho' c_\infty^2 = \frac{\partial^2 \phi_{ij}}{\partial x_i \partial x_j} = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \int_{-T}^T T_{ij}(\mathbf{y}, \tau) G(\mathbf{x}, t | \mathbf{y}, \tau) dV d\tau$$

Lighthill's Equation

$$\rho' c_{\infty}^2(\mathbf{x}, t) = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \frac{1}{4\pi |\mathbf{x} - \mathbf{y}|} [T_{ij}(\mathbf{y}, \tau)]_{\tau=\tau^*} dV$$

$$\tau^* = t - \frac{|\mathbf{x} - \mathbf{y}|}{c_{\infty}}$$



The sound field is now the double divergence of the propagating waves, not the source term

Lighthill's Equation in Far Field

$$\rho' c_\infty^2(\mathbf{x}, t) = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \frac{1}{4\pi|\mathbf{x} - \mathbf{y}|} [T_{ij}(\mathbf{y}, \tau)]_{\tau=\tau^*} dV$$

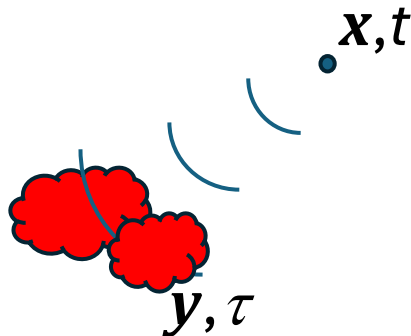
$$\rho' c_\infty^2(\mathbf{x}, t) \approx \frac{1}{4\pi|\mathbf{x}|} \frac{\partial^2}{\partial x_i \partial x_j} \int_V [T_{ij}(\mathbf{y}, \tau)]_{\tau=\tau^*} dV$$

$$\frac{\partial^2 f(\tau^*)}{\partial x_i \partial x_j} \approx \frac{\partial \tau^*}{\partial x_i} \frac{\partial \tau^*}{\partial x_j} \frac{\partial^2 f(\tau^*)}{\partial \tau^{*2}}$$

$$\tau^* = t - \frac{|\mathbf{x} - \mathbf{y}|}{c_\infty}$$

$$\frac{\partial \tau^*}{\partial x_i} \approx \frac{1}{c_\infty} \frac{x_i}{|\mathbf{x}|}$$

Turbulence



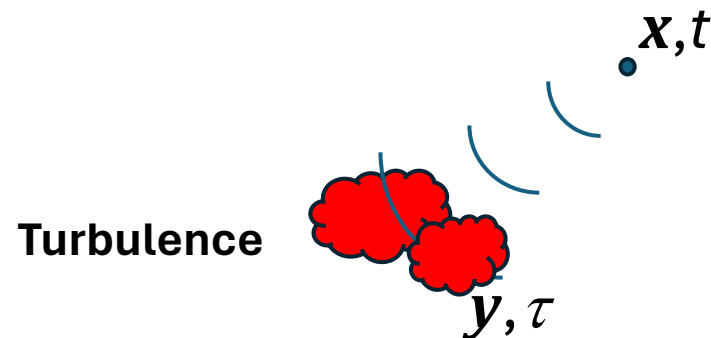
*The rate of change of the
wave motion is much faster
than the decay with
distance*

Lighthill's Equation in the Far Field

$$\rho' c_\infty^2(\mathbf{x}, t) = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \frac{1}{4\pi |\mathbf{x} - \mathbf{y}|} [T_{ij}(\mathbf{y}, \tau)]_{\tau=\tau^*} dV$$

$$\rho' c_\infty^2(\mathbf{x}, t) \approx \frac{x_i x_j}{4\pi |\mathbf{x}|^3 c_\infty^2} \int_V \left[\frac{\partial^2 T_{ij}(\mathbf{y}, \tau)}{\partial \tau^2} \right]_{\tau=\tau^*} dV$$

$$\tau^* = t - \frac{|\mathbf{x} - \mathbf{y}|}{c_\infty}$$



*Note the quadrupole type
directionality and the
speed of sound
dependence*

Scaling on Flow Speed

$$\rho' c_{\infty}^2(\mathbf{x}, t) \approx \frac{x_i x_j}{4\pi |\mathbf{x}|^3 c_{\infty}^2} \int_V \left[\frac{\partial^2 T_{ij}(\mathbf{y}, \tau)}{\partial \tau^2} \right]_{\tau=\tau^*} dV$$

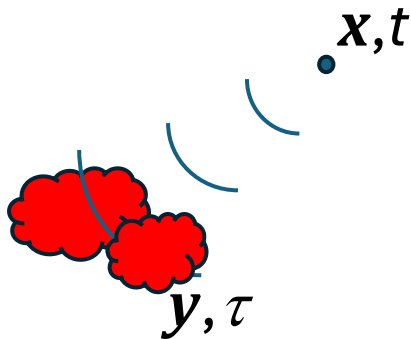
$$T_{ij} = \rho v_i v_j \sim \rho_o U^2$$

$$\tau \sim L/U$$

$$\rho' c_{\infty}^2(\mathbf{x}, t) = p \sim \frac{\rho_o x_i x_j U^4 V}{4\pi |\mathbf{x}|^3 c_{\infty}^2 L^2}$$

$$p^2 \sim \left(\frac{\rho_o x_i x_j V}{4\pi |\mathbf{x}|^3 c_{\infty}^2 L^2} \right)^2 U^8$$

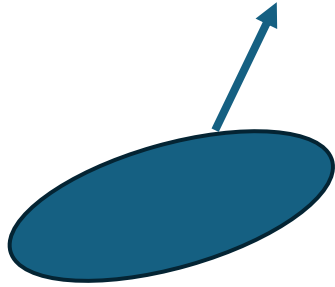
Turbulence



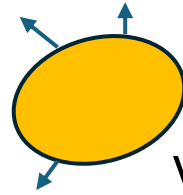
- The mean square pressure, or intensity, scales with the Eighth power of the flow speed
- Halving the flow speed gives a 24dB noise reduction

Adding Surfaces to the Flow

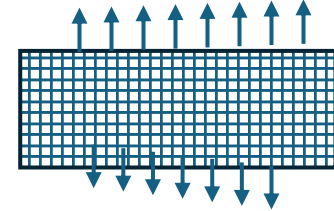
The Ffowcs Williams and Hawkings Equation



Solid body in motion

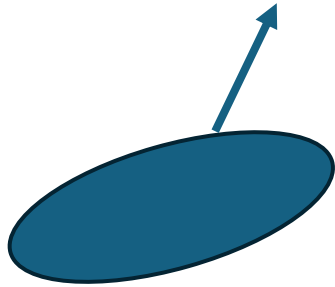


Vibrating surface

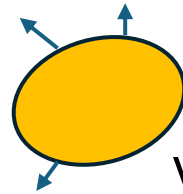


Permeable surface
such as a CFD grid

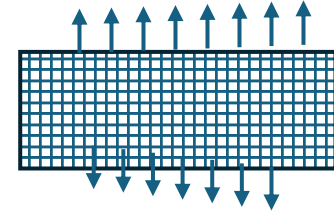
The Ffowcs Williams and Hawkings Equation



Solid body in motion

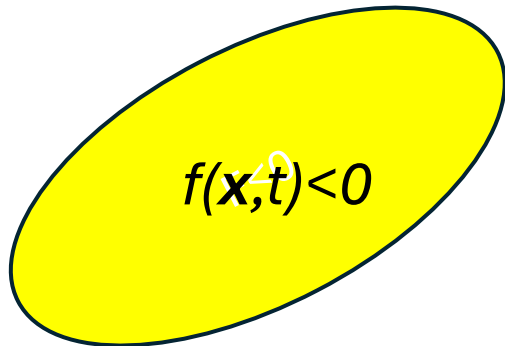


Vibrating surface



Permeable surface
such as a CFD grid

$$f(\mathbf{x},t) > 0$$



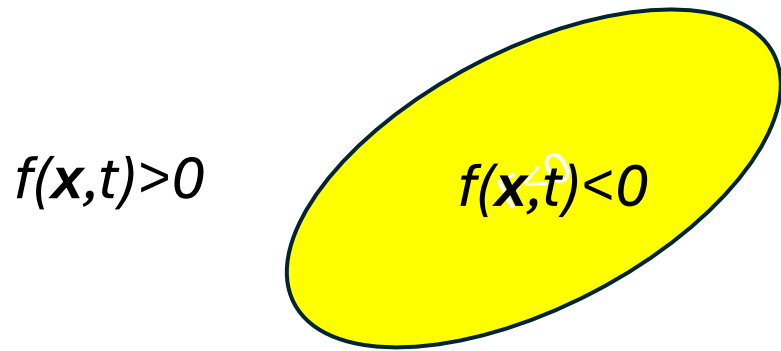
$$f(\mathbf{x},t) < 0$$

The surface is defined as a function of space and time by the function $f(\mathbf{x},t)=0$

$$f(\mathbf{x},t) = x_1^2 + x_2^2 + (x_3 - Ut)^2 - a^2$$

sphere moving at constant speed

The Ffowcs Williams and Hawkings Equation



Define new variables

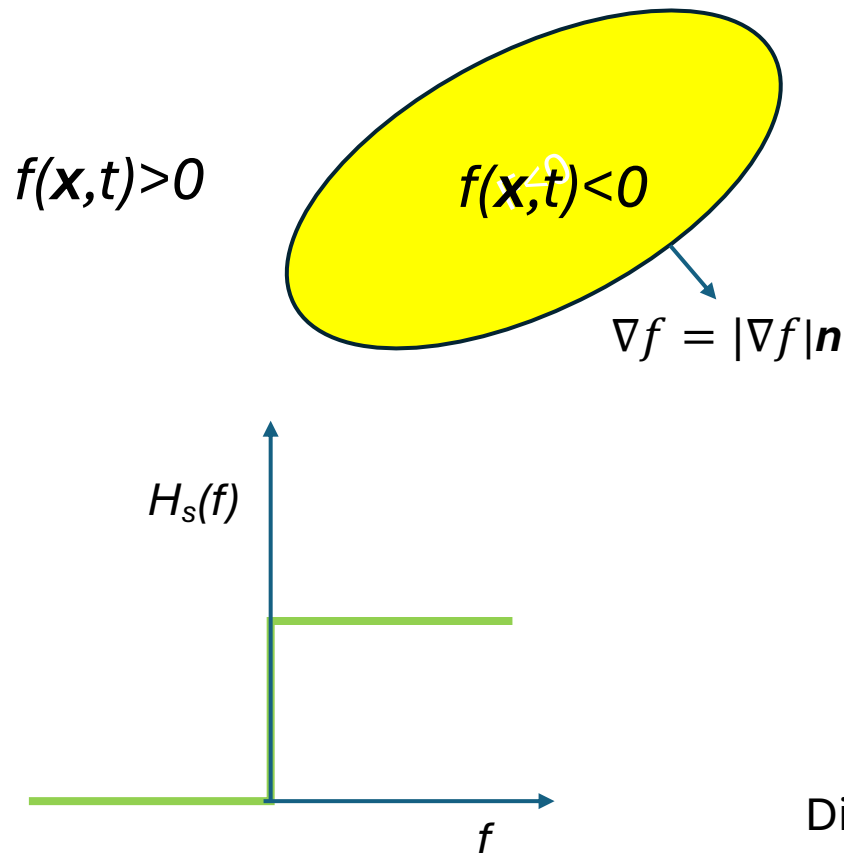
$$vH_s(f) \quad pH_s(f) \quad \rho H_s(f)$$

$$H_s(f) = 0. \quad \text{inside surface}$$

$$H_s(f) = 1. \quad \text{outside surface}$$

Each new variable is zero inside the surface but equal to its defined value outside the surface

The Ffowcs Williams and Hawkings Equation



Operators applied to new variables

$$\nabla \cdot (\mathbf{v} H_s(f)) = H_s(f) \nabla \cdot (\mathbf{v}) + \mathbf{v} \cdot \nabla H_s(f)$$

$$\nabla H_s(f) = \frac{\partial H_s(f)}{\partial f} \nabla f = \delta(f) \nabla f$$

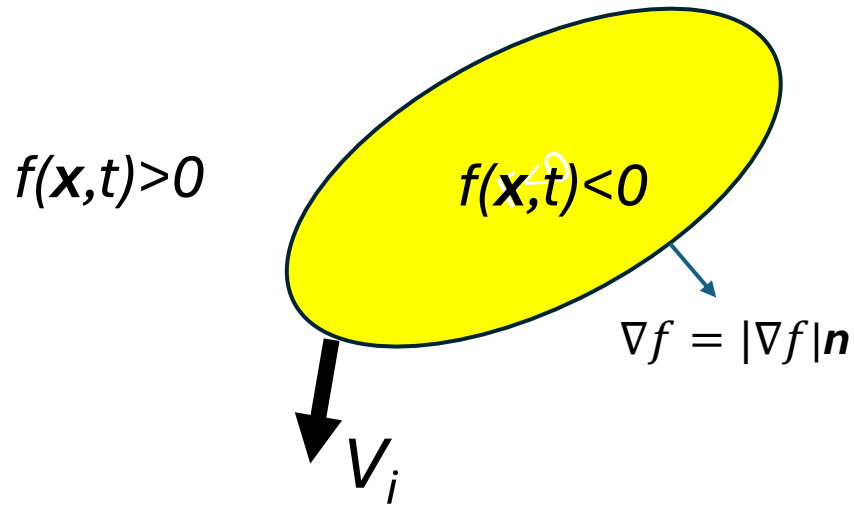
$$\nabla \cdot (\mathbf{v} H_s(f)) = H_s(f) \nabla \cdot (\mathbf{v}) + \mathbf{v} \cdot \mathbf{n} |\nabla f| \delta(f)$$

Divergence of variable outside surface
but is zero inside the surface

Velocity normal to surface

The Ffowcs Williams and Hawkings Equation

New Continuity equation



$$\frac{\partial(\rho' H_s)}{\partial t} + \nabla \cdot (\rho \mathbf{v} H_s) = H_s \left(\frac{\partial \rho'}{\partial t} + \nabla \cdot (\rho \mathbf{v}) \right) + (\rho \mathbf{v} - \rho' \mathbf{V}) \cdot \mathbf{n} |\nabla f| \delta(f)$$

zero

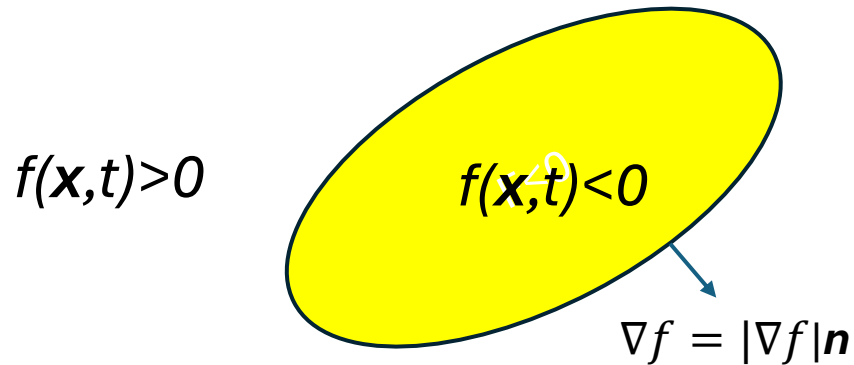
$$\frac{\partial(\rho' H_s)}{\partial t} + \frac{\partial}{\partial x_i} (\rho v_i H_s) = (\rho v_i - \rho' V_i) \cdot n_i |\nabla f| \delta(f)$$

Continuity equation

$$\frac{\partial(\rho v_i H_s)}{\partial t} + \frac{\partial}{\partial x_j} (\rho v_i v_j H_s + p_{ij} H_s) = (\rho v_i (v_j - V_j) + p_{ij}) n_j |\nabla f| \delta(f)$$

Momentum equation

The Ffowcs Williams and Hawkings Equation



New Wave Equation

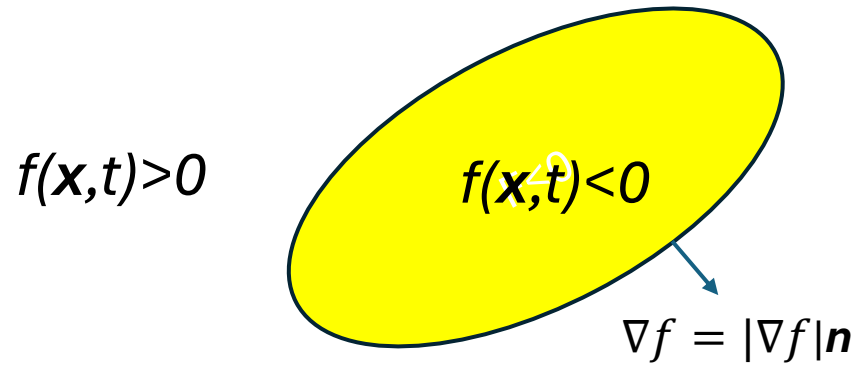
$$\frac{\partial^2(\rho' H_s)}{\partial t^2} - c_\infty^2 \frac{\partial^2(\rho' H_s)}{\partial x_i^2} = \frac{\partial^2(T_{ij} H_s)}{\partial x_i \partial x_j} - \frac{\partial}{\partial x_i} ((\rho v_i (v_j - V_j) + p_{ij}) n_j |\nabla f| \delta(f)) + \frac{\partial}{\partial t} ((\rho v_i - \rho' V_i) \cdot n_i |\nabla f| \delta(f))$$

Surface momentum flux and pressure stress

Surface Motion Term

Lighthill's Source

The solution to the FW-H Equation



$$G(\mathbf{x}, t | \mathbf{y}, \tau) = \delta \left(t - \tau - \frac{|\mathbf{x} - \mathbf{y}|}{c_\infty} \right) / 4\pi |\mathbf{x} - \mathbf{y}|$$

$$\rho' c_\infty^2 H_s = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \int_{-T}^T T_{ij} H_s G dV d\tau$$

Lighthill's Source

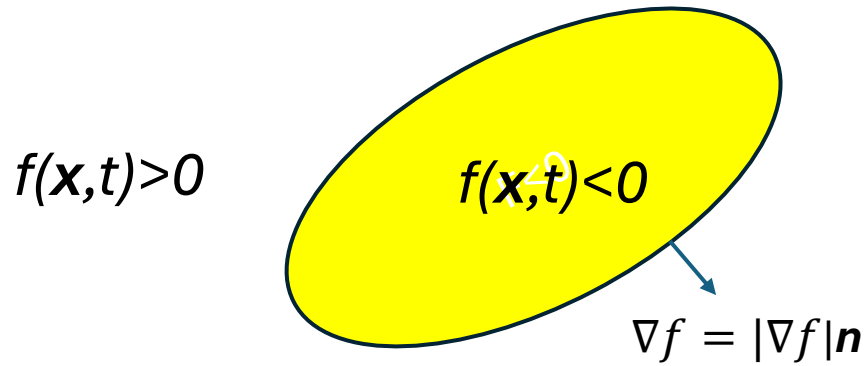
$$- \frac{\partial}{\partial x_i} \int_V \int_{-T}^T ((\rho v_i (v_j - V_j) + p_{ij}) n_j |\nabla f| \delta(f)) G dV d\tau$$

Surface momentum flux and pressure stress

$$+ \frac{\partial}{\partial t} \int_V \int_{-T}^T ((\rho v_i - \rho' V_i) \cdot n_i |\nabla f| \delta(f)) G dV d\tau$$

Surface Motion Term

The solution to the FW-H Equation



$$G(\mathbf{x}, t | \mathbf{y}, \tau) = \frac{\delta\left(t - \tau - \frac{r(\tau)}{c_\infty}\right)}{4\pi r(\tau)}$$

$$r(\tau) = |\mathbf{x} - \mathbf{y}(\tau)|$$

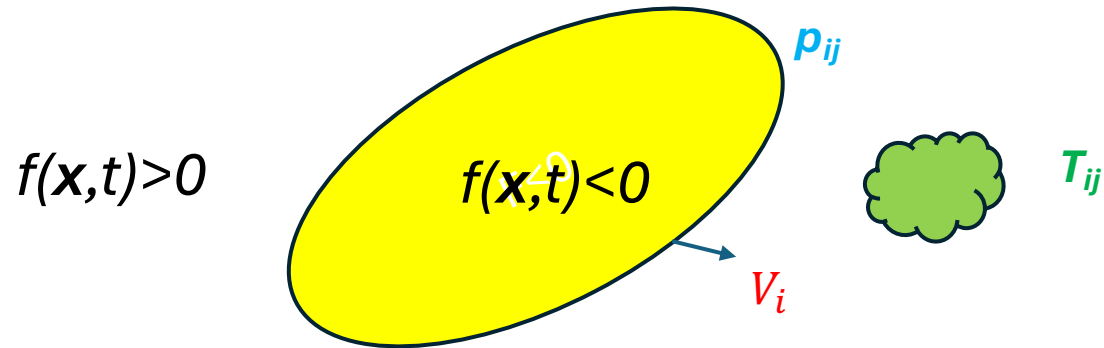
$$\int_V \int_{-T}^T q_j(\mathbf{y}, \tau) n_j |\nabla f| \delta(f) G(\mathbf{x}, t | \mathbf{y}, \tau) dV d\tau$$

$$\frac{1}{c_\infty} \frac{\partial r}{\partial \tau} = -M_r$$

$$= \int_S \int_{-T}^T q_j(\mathbf{y}, \tau) n_j G(\mathbf{x}, t | \mathbf{y}, \tau) dS d\tau$$

$$= \int_S \left[\frac{q_j(\mathbf{y}, \tau) n_j}{4\pi r(\tau)} \right]_{\tau=\tau^*} dS$$

The solution to the FW-H Equation

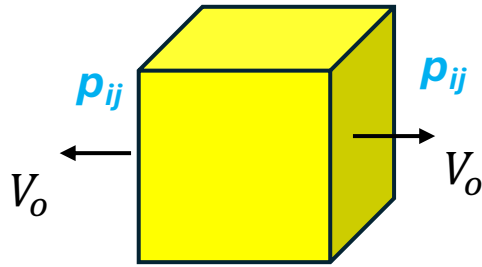


$$\rho' c_{\infty}^2 H_s = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \left[\frac{T_{ij} H_s}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dV \quad \text{Lighthill's Source}$$

$$- \frac{\partial}{\partial x_i} \int_S \left[\frac{(\rho v_i (v_j - V_j) + p_{ij}) n_j}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dS \quad \text{Surface momentum flux and pressure stress}$$

$$+ \frac{\partial}{\partial t} \int_S \left[\frac{(\rho v_i - \rho' V_i) \cdot n_i}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dS \quad \text{Surface Motion Term}$$

Some Examples- Cube with 2 sides moving



The surface is impenetrable so $\mathbf{v}=\mathbf{V}$

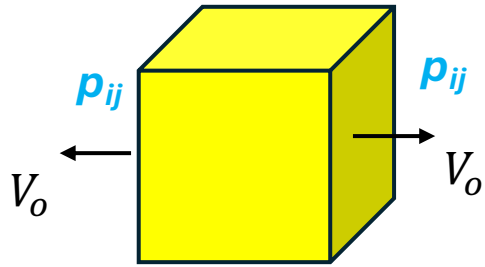
$$(\rho v_i - \rho' V_i) n_i = (\rho v_i - (\rho - \rho_\infty) V_i) n_i = \rho_\infty V_i n_i$$

$$\rho' c_\infty^2 H_s = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \left[\frac{T_{ij} H_s}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dV \quad \text{Lighthill's Source}$$

$$- \frac{\partial}{\partial x_i} \int_S \left[\frac{(\cancel{\rho v_i (v_j - V_j)} + p_{ij}) n_j}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dS \quad \text{Surface momentum flux and pressure stress}$$

$$+ \frac{\partial}{\partial t} \int_S \left[\frac{(\rho v_i - \rho' V_i) \cdot n_i}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dS \quad \text{Surface Motion Term}$$

Some Examples- Cube with 2 sides moving



- Assume observer is in the far field so $r=|\mathbf{x}|$
- Ignore the non linear terms
- Assume $V_o \ll c_\infty$

~~$$\rho' c_\infty^2 H_s = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \left[\frac{T_{ij} H_s}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dV$$~~

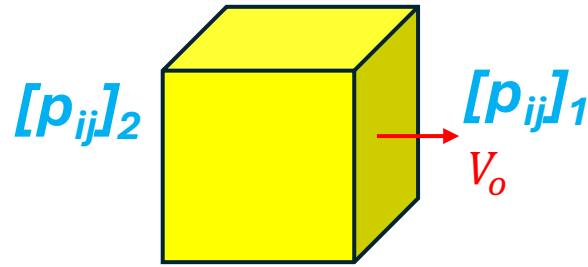
$$- \frac{1}{4\pi |\mathbf{x}|} \frac{\partial}{\partial x_i} \int_S [p_{ij} n_j]_{\tau=\tau^*} dS \quad \text{Net Surface Force}$$

$$+ \frac{1}{4\pi |\mathbf{x}|} \frac{\partial}{\partial t} \int_S [\rho_\infty V_i \cdot n_i]_{\tau=\tau^*} dS$$

**The net force is zero
so only the volume
displacement term
remains**

Surface Motion Term

Some Examples- Cube with One side moving



- Assume observer is in the far field so $r=|\mathbf{x}|$
- Ignore the non linear terms
- Assume $V_o \ll c_\infty$

$$\rho' c_\infty^2 H_S =$$

$$- \frac{1}{4\pi|\mathbf{x}|} \frac{\partial}{\partial x_i} \int_S [p_{ij} n_j]_{\tau=\tau^*} dS$$

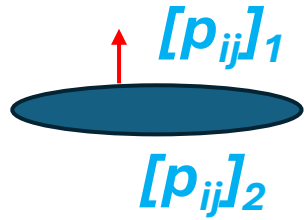
Net Surface Force

$$+ \frac{1}{4\pi|\mathbf{x}|} \frac{\partial}{\partial t} \int_S [\rho_\infty V_i \cdot n_i]_{\tau=\tau^*} dS$$

The net force is not zero so both terms are needed

Surface Motion Term

Some Examples- Stationary Airfoil with unsteady load



- Assume observer is in the far field so $r=|\mathbf{x}|$
- Ignore the non linear terms
- Assume $V_o=0$

$$\rho' c_\infty^2 H_s = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \left[\frac{T_{ij} H_s}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dV$$

The net force is the only term remaining

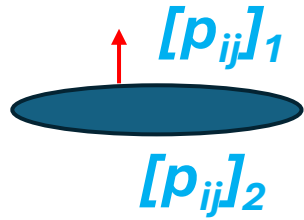
$$- \frac{1}{4\pi |\mathbf{x}|} \frac{\partial}{\partial x_i} \int_S [p_{ij} n_j]_{\tau=\tau^*} dS$$

Net Surface Force

$$+ \frac{1}{4\pi |\mathbf{x}|} \frac{\partial}{\partial t} \int_S [\rho_\infty V_i \cdot n_i]_{\tau=\tau^*} dS$$

Surface Motion Term

Some Examples- Stationary Airfoil with unsteady load



- Assume observer is in the far field so $r=|\mathbf{x}|$
- Ignore the non linear terms
- Assume $V_o=0$

Pressure integrated at the correct retarded time

$$-\frac{1}{4\pi|\mathbf{x}|}\frac{\partial}{\partial x_i}\int_S [p_{ij}n_j]_{\tau=\tau^*}dS = \frac{x_i}{4\pi|\mathbf{x}|^2c_\infty}\frac{\partial}{\partial t}\int_S [p_{ij}n_j]_{\tau=\tau^*}dS$$

$$= \frac{x_i}{4\pi|\mathbf{x}|^2c_\infty}\left[\frac{\partial F_i(\tau)}{\partial \tau}\right]_{\tau=\tau^*}$$

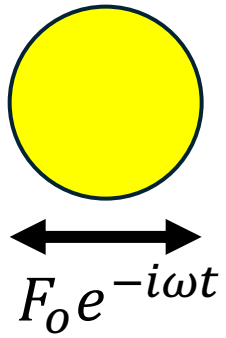
Net Surface Force if the surface is compact

$$p \sim \rho_\infty U^3 A / L c_\infty r$$

$$F_i \sim \rho_\infty U^2 A \quad t \sim L/U.$$

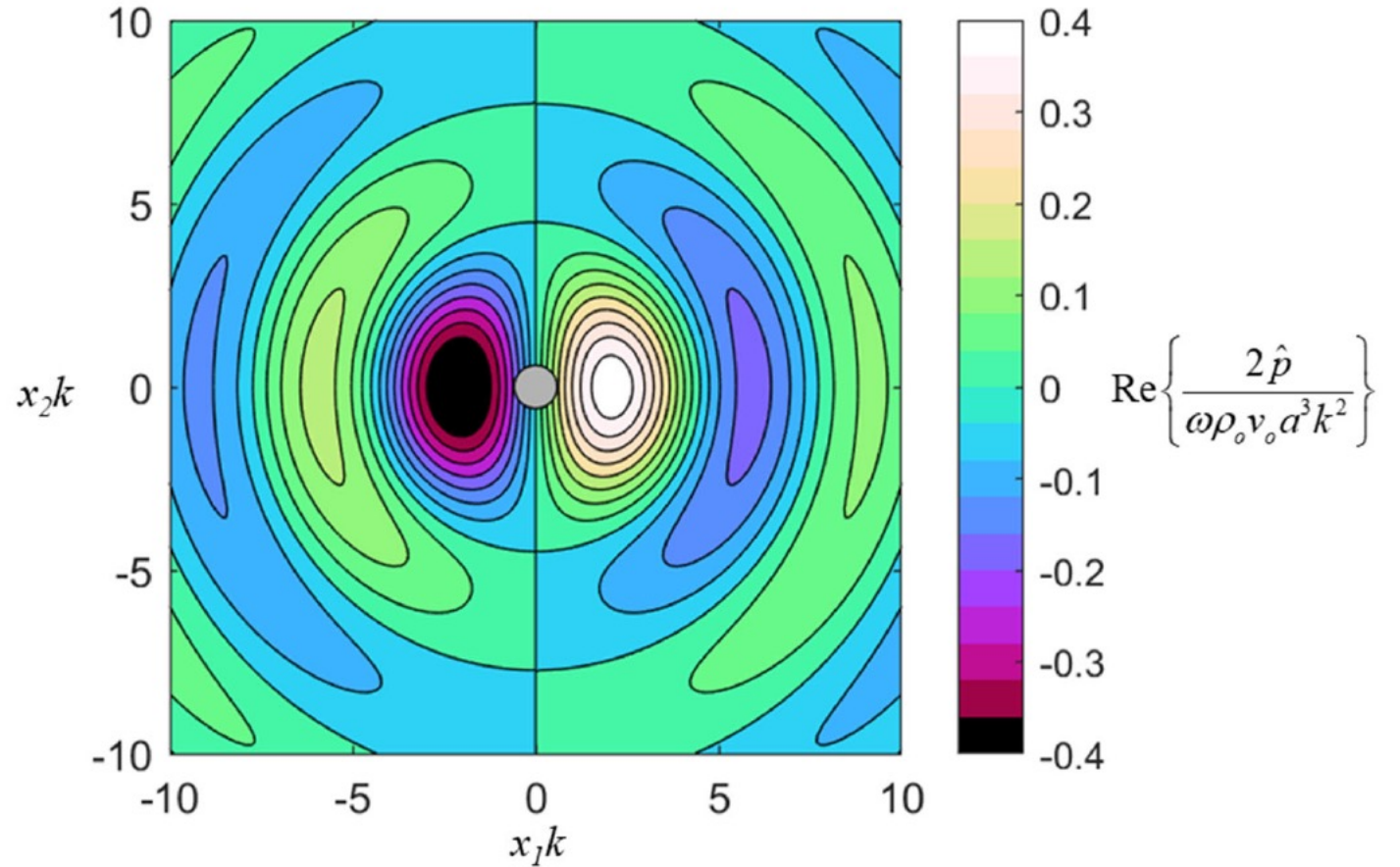
$$p^2 \sim U^6$$

Cosine Directionality



$$\frac{x_1}{r} = \cos\theta$$

$$\hat{p}(r) = -ik \cos\theta e^{ikr} \left(\frac{\hat{F}}{4\pi r} \right) \left(1 - \frac{1}{ikr} \right)$$



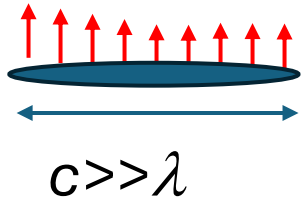
Loading Noise Scales as U^6

$$p^2 \sim U^6$$

- Halving the flow speed gives an 18 dB noise reduction for no change in shape
- Since lift $\sim U^2 A$ and noise as $\sim U^6 A^2 = U^2 (U^2 A)^2$ halving speed at constant lift gives a 6dB noise reduction
- For a propeller thrust $\sim U_{tip}^2 R^2$ and noise $\sim U_{tip}^6 R^4 = (U_{tip}^2 R^2)^3 / R^2$ so noise is reduced by increasing the diameter

Some Examples- Large Flat Surface

$\Delta p = \text{pressure jump}$



- Assume observer is in the far field so $r = |\mathbf{x}| \gg c \gg \lambda$
- Ignore the non linear terms
- Assume $V_o = 0$

$$-\frac{1}{4\pi|\mathbf{x}|} \frac{\partial}{\partial x_i} \int_S [p_{ij} n_j]_{\tau=\tau^*} dS = \frac{x_1}{4\pi|\mathbf{x}|^2 c_\infty} \frac{\partial}{\partial t} \int_\Sigma [\Delta p]_{\tau=\tau^*} d\Sigma$$

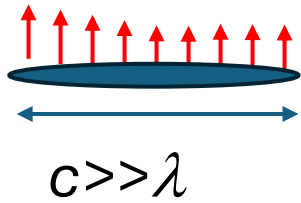
$$\Delta p(\mathbf{y}, \tau) = P_o(y_1, y_3) \exp(-i\omega t)$$

Far Field Approximation for the correct retarded time

$$\frac{-i\omega x_1}{4\pi|\mathbf{x}|^2 c_\infty} \int_0^c \int_{-\frac{b}{2}}^{\frac{b}{2}} P_o(y_1, y_3) e^{-i\omega t - ik|\mathbf{x}| + ik\mathbf{x} \cdot \mathbf{y}/|\mathbf{x}|} dy_1 dy_3$$


Some Examples- Large Flat Surface

$\Delta p = \text{pressure jump}$



- Assume observer is in the far field so $r = |\mathbf{x}| \gg c \gg \lambda$
- Ignore the non linear terms
- Assume $V_o = 0$

$$P_o(y_1, y_3) = P_o = \text{const}$$

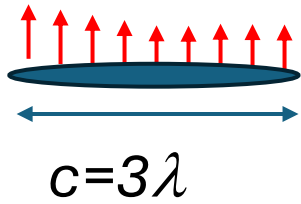
$$\frac{-i\omega x_1 P_o e^{-i\omega t - ik|x|}}{4\pi |x|^2 c_\infty} \int_0^c \int_{-\frac{b}{2}}^{\frac{b}{2}} e^{ik(x_1 y_1 + x_3 y_3)/|x|} dy_1 dy_3$$

$$= \frac{-i\omega x_1 b c P_o e^{-i\omega t - ik|x|}}{4\pi |x|^2 c_\infty} \left\{ \frac{\sin\left(\frac{kx_1 c}{2|x|}\right)}{\frac{kx_1 c}{2|x|}} \right\} \left\{ \frac{\sin\left(\frac{kx_3 b}{2|x|}\right)}{\frac{kx_3 b}{2|x|}} \right\}$$

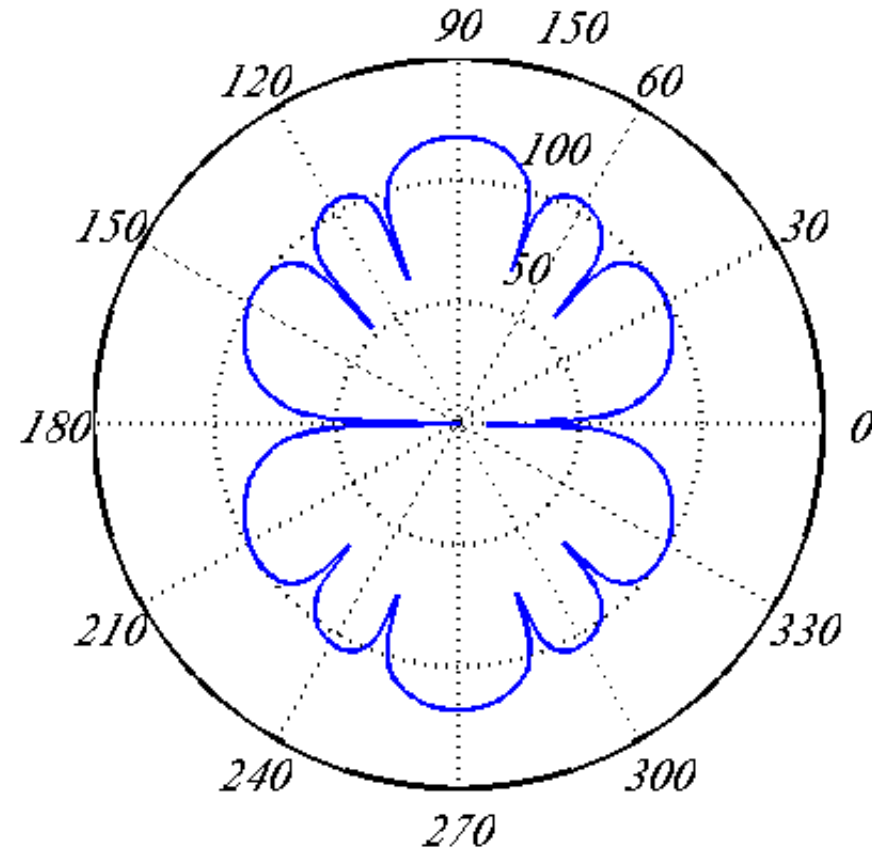
Additional Directionality Terms

Some Examples- Large Flat Surface

$\Delta p = \text{pressure jump}$

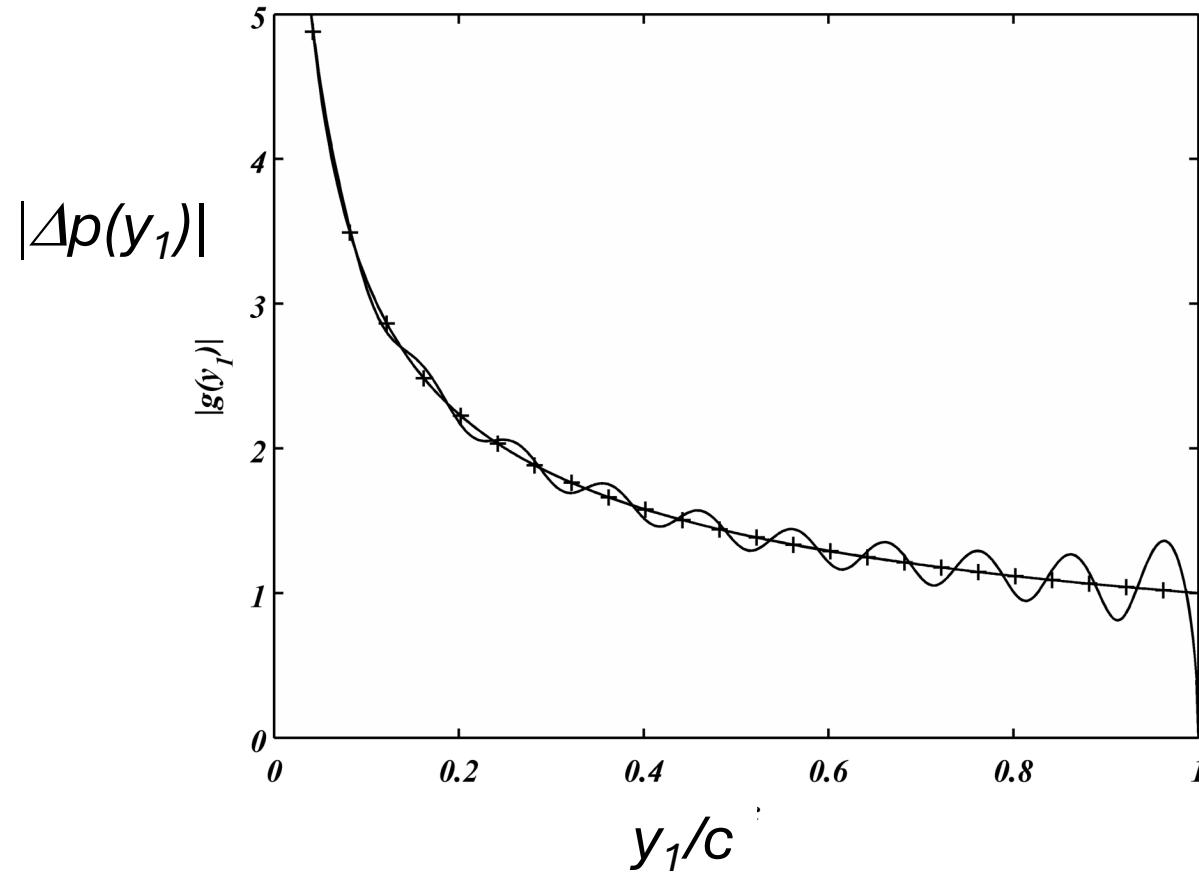


$$P_o(y_1, y_3) = P_o = \text{const}$$



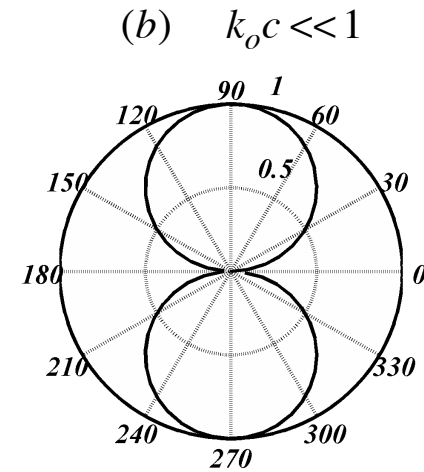
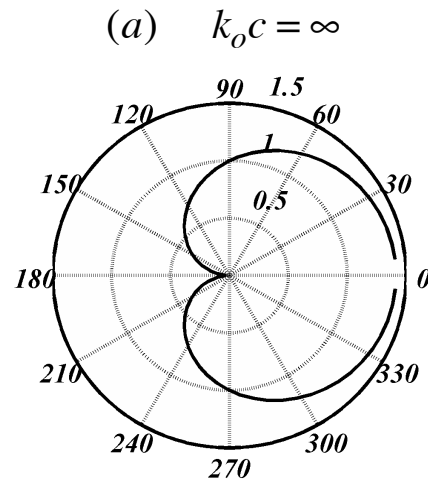
$$\frac{-i\omega x_1 b c P_o e^{-i\omega t - ik|x|}}{4\pi|x|^2 c_\infty} \left\{ \frac{\sin\left(\frac{kx_1 c}{2|x|}\right)}{\frac{kx_1 c}{2|x|}} \right\} \left\{ \frac{\sin\left(\frac{kx_3 b}{2|x|}\right)}{\frac{kx_3 b}{2|x|}} \right\}$$

Unsteady Loading Distribution on Airfoil

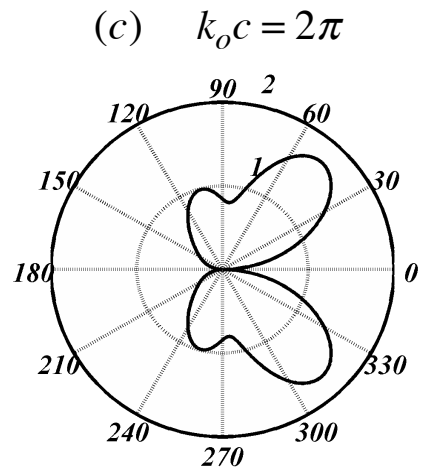


Far Field Directionality

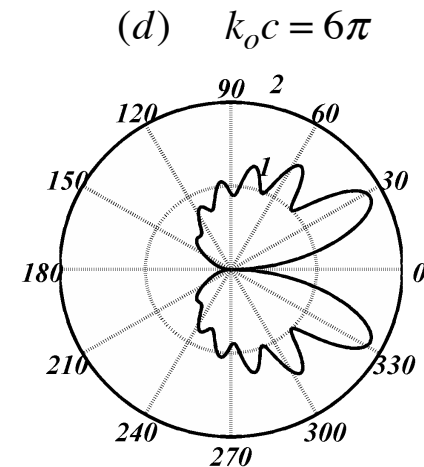
Semi-Infinite
Blade Chord



Chord $\ll \lambda$



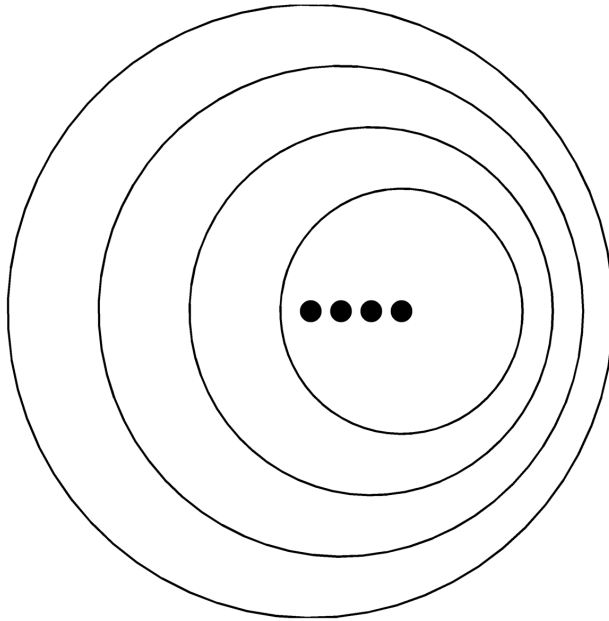
Chord $= \lambda$



Chord $= 3\lambda$

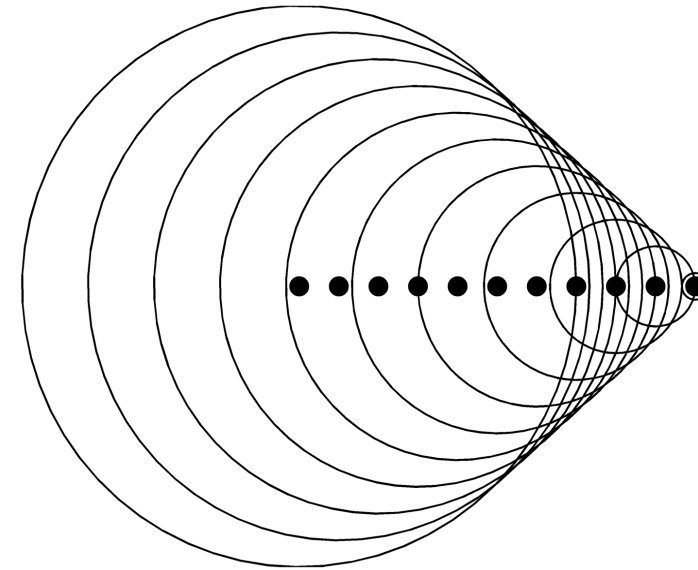
Moving Sources

Moving Sources



Subsonic Motion

(a)



Supersonic Motion

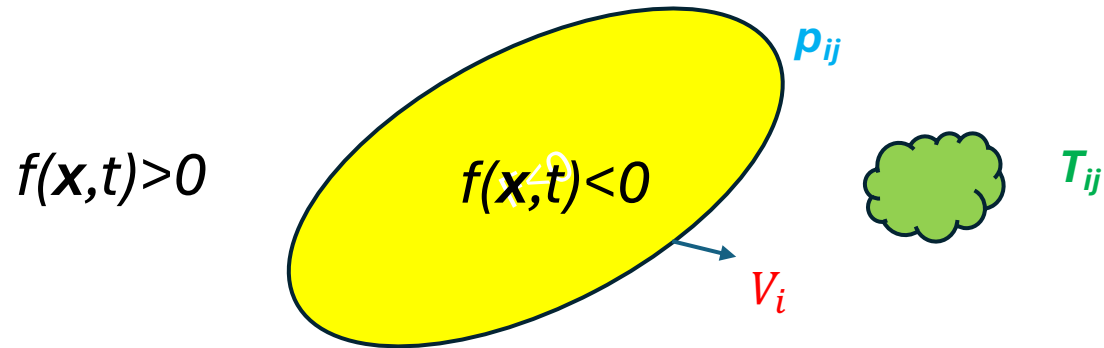
(b)

$$t = \tau + |\mathbf{x} - \mathbf{y}(\tau)|/c_\infty$$

$$\frac{\partial f(\tau)}{\partial t} = \frac{\partial \tau}{\partial t} \frac{\partial f}{\partial \tau} = \frac{1}{1 - M_r} \frac{\partial f}{\partial \tau}$$

$$M_r = V \cos \theta / c_\infty$$

The solution to the FW-H Equation

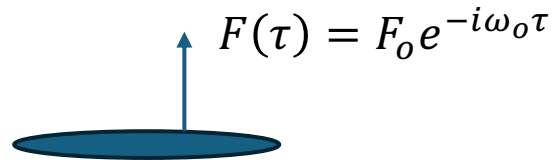


$$\rho' c_\infty^2 H_s = \frac{\partial^2}{\partial x_i \partial x_j} \int_V \left[\frac{T_{ij} H_s}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dV \quad \text{Lighthill's Source}$$

$$- \frac{\partial}{\partial x_i} \int_S \left[\frac{p_{ij} n_j}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dS \quad \text{Surface momentum flux and pressure stress}$$

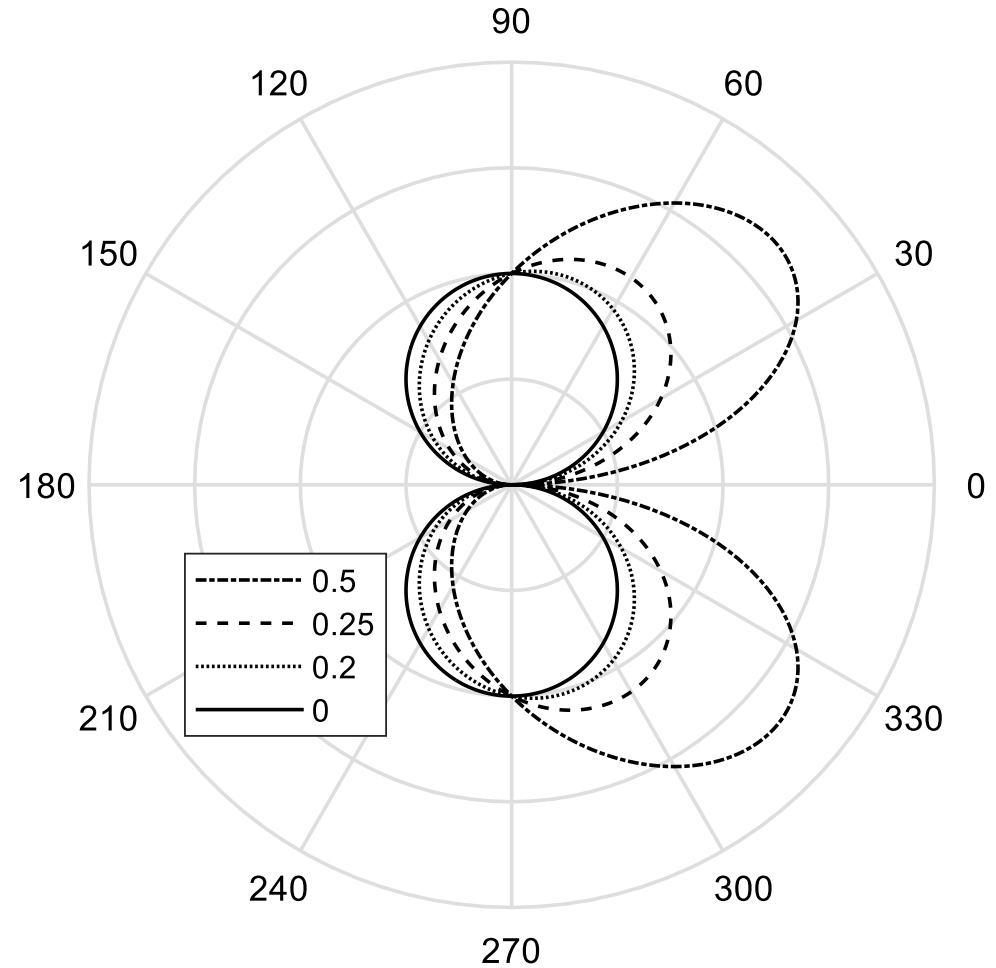
$$+ \frac{\partial}{\partial t} \int_S \left[\frac{\rho_\infty V_i \cdot n_i}{4\pi r(\tau) |1 - M_r|} \right]_{\tau=\tau^*} dS \quad \text{Surface Motion Term}$$

Dipole in Motion



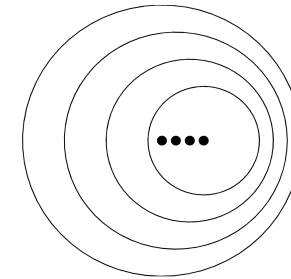
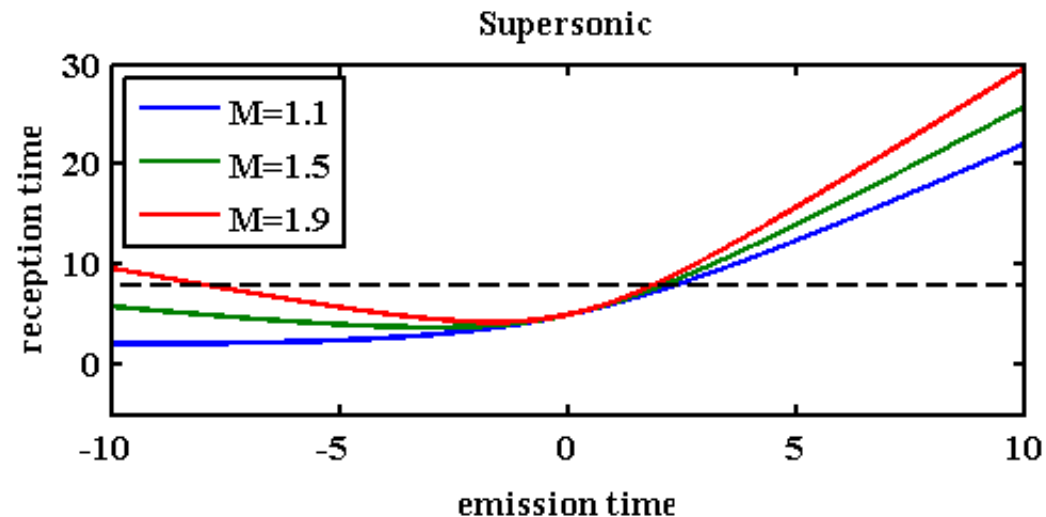
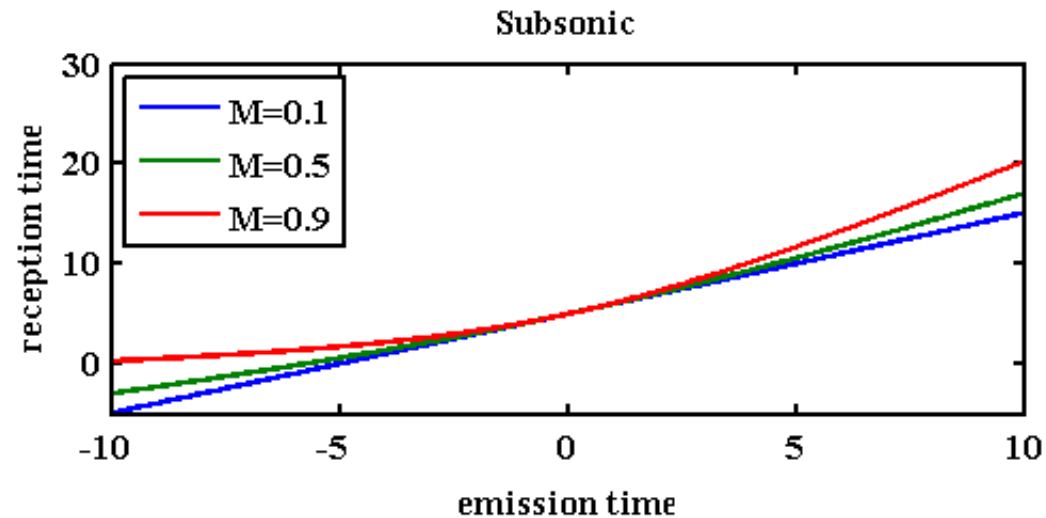
In the far field for a harmonic source

$$\rho' c_{\infty}^2 H_s = \left[\frac{-i \omega_o x_2 F_o e^{-i \omega_o \left(t - \frac{|x|}{c_{\infty}} \right) / (1 - M_r)}}{4 \pi |x|^2 (1 - M_r)^2} \right]$$

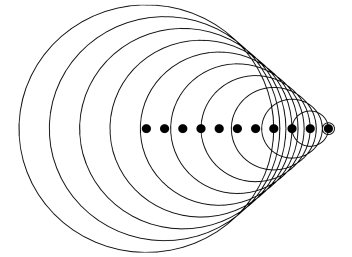


Note the Doppler amplification caused by the source motion

Reception and Emission Time



(a)

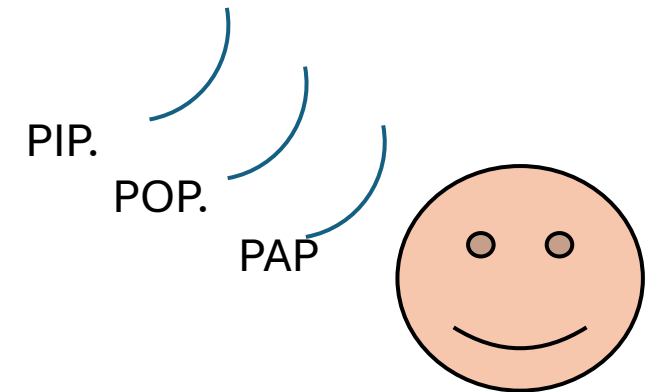
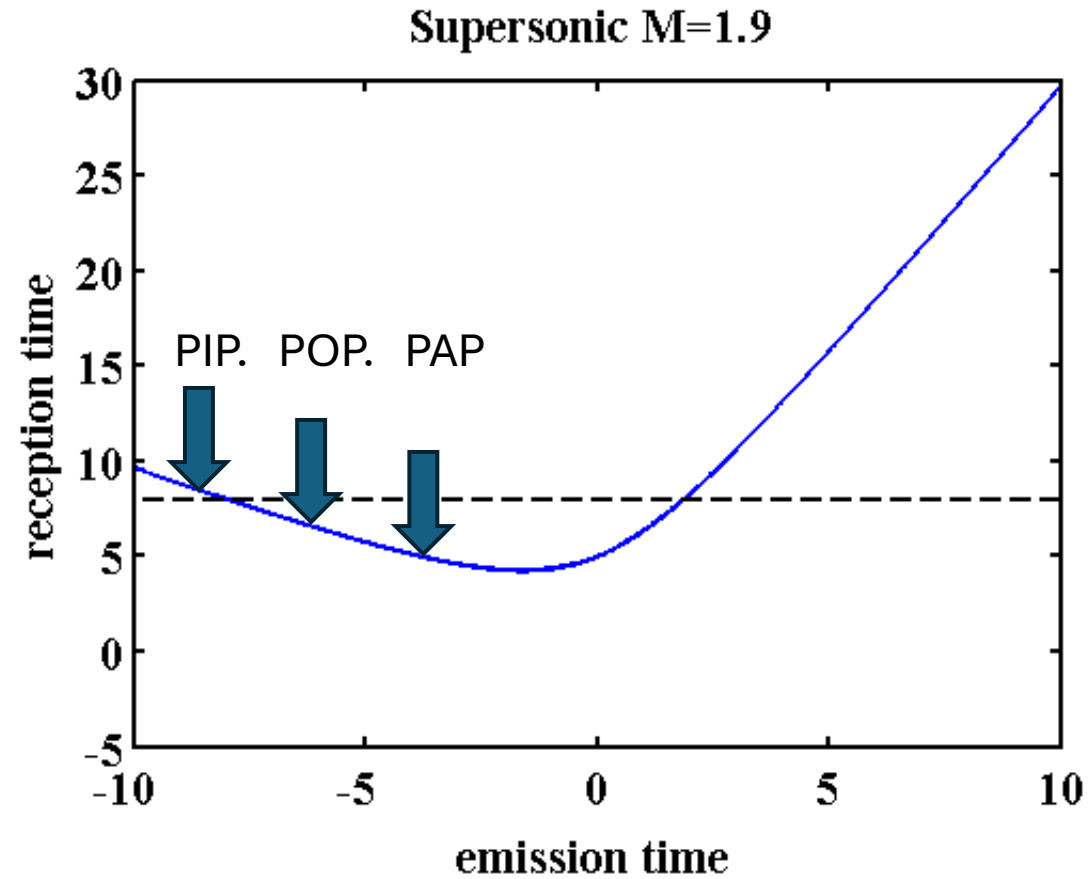


(b)

For subsonic source there is only one emission time for each reception time

For supersonic source there can be two emission times for the same reception time

PIP POP PAP. (Lighthill)



The observer hears the sound signature in reverse